PHY 835: Machine Learning in Physics Lecture 16: Decision Trees & kNN March 14, 2024



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### Outline for today

- Many ML models are designed to solve classification or regression problems.
- Simple Classifiers:
  - Decision Trees
  - kNN: Finding neighbors

• Reference: "Machine Learning for Physics and Astronomy" by Viviana Acquaviva, Princeton University Press, Chapter 2.

# We can now design a decision tree in a (astro)physics context.

#### OUR FIRST EXAMPLE WILL BE A SUPERVISED CLASSIFICATION PROBLEM, IN WHICH WE ATTEMPT TO PREDICT IF A PLANET IS HABITABLE, BASED ON ITS DISTANCE FROM PARENT STAR, MASS OF PARENT STAR, AND ORBITAL PERIOD.

### Search for Habitable Planets

- The problem: search for intelligent life beyond Earth.
- More than 5000 exoplanets: <u>https://exoplanets.nasa.gov/alien-worlds/</u> <u>historic-timeline/#first-transiting-exoplanet-observed</u>
- Finding habitable planets (density and temperature conditions etc that are compatible with the development of life).
- Planet Habitability Laboratory website: <u>https://phl.upr.edu/projects/</u> <u>habitable-exoplanets-catalog</u> contains data for thousands of planets and collects a variety of features.
- We consider a learning set composed of 18 instances and their 3 features.

#### DATA FROM THE PLANET HABITABILITY LAB AT ARECIBO OBSERVATORY



NAME	Stellar Mass (M <sub>☉</sub> )	Orbital Period (days)	Distance (AU)	Habitable?
Kepler-736 b	0.86	3.60	0.0437	0
Kepler-636 b	0.85	16.08	0.1180	0
Kepler-887 c	1.19	7.64	0.0804	0
Kepler-442 b	0.61	112.30	0.4093	1
Kepler-772 b	0.98	12.99	0.1074	0
Teegarden's Star b	0.09	4.91	0.0252	1
K2-116 b	0.69	4.66	0.0481	0
GJ 1061 c	0.12	6.69	0.035	1
HD 68402 b	1.12	1103	2.1810	0
Kepler-1544 b	0.81	168.81	0.5571	1
Kepler-296 e	0.5	34.14	0.1782	1
Kepler-705 b	0.53	56.06	0.2319	1
Kepler-445 c	0.18	4.87	0.0317	0
HD 104067 b	0.62	55.81	0.26	0
GJ 4276 b	0.41	13.35	0.0876	0
Kepler-296 f	0.5	63.34	0.2689	1
Kepler-63 b	0.98	9.43	0.0881	0
GJ 3293 d	0.42	48.13	0.1953	1

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outlier	Kepler-736 b	0.86	3.60	0.0437	0
	Kepler-636 b	0.85	16.08	0.1180	0
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### Seems reasonable from Kepler's laws

-	NAME	Stellar Mass $(M_{\odot})$	Orbital Period (days)	Distan	e (AU)	Habitable?
-	Kepler-736 b	0.86	3.60	0.0	437	0
	Kepler-636 b	0.85	16.08	0.1	180	0
	Kepler-887 c	1.19	7.64	0.0	804	0
-	Kepler-442 b	0.61	112.30	0.4	093	1
-	Kepler-772 b	0.98	12.99	0.1	074	0
-	Teegarden's Star b	0.09	4.91	0.0	252	1
-	K2-116 b	0.69	4.66	0.0	481	0
outlier	GJ 1061 c	0.12	6.69	9.0	35	1
	HD 68402 b	1.12	1103	2.18	810	0
	Kepler-1544 b	0.81	168.81	0.53	571	1
	Kepler-296 e	0.5	34.14	0.1	782	1
	Kepler-705 b	0.53	56.06	0.23	319	1
	Kepler-445 c	0.18	4.87	0.0	317	0
	HD 104067 b	0.62	55.81	0.5	26	0
	GJ 4276 b	0.41	13.35	0.0	876	0
	Kepler-296 f	0.5	63.34	0.2	689	1
	Kepler-63 b	0.98	9.43	0.0	881	0
	GJ 3293 d	0.42	48.13	0.19	953	1

### Comments on the dataset

- Well balanced: 10 examples in one category (not habitable) and 8 in the other (habitable).
- A factor that determines whether a planet is habitable is temperature, which likely depends on the energy it receives from its parent star.
- The planet's temperature therefore depends on the star's luminosity and its **distance** from the planet.
- Mass of a star is a decent tracer of its luminosity (as is the case for main sequence stars).
- Reality is more complicated: the energy budget of each planet also depends on other features, e.g., properties of its atmosphere, & whether it has an internal energy source.
- Mass/luminosity relationship is monotonic only for main sequence stars, which make up only about 90% of the total.

### Predicting planet habitability

- On canvas, under DT-kNN-Notebooks, you will find a Jupyter Notebook: Intro\_DT\_HabPlanets.ipynb and a dataset: HPLearningSet.csv.
- Split into a training and a test set:

```
LearningSet = pd.read_csv (`HPLearningSet.csv')
TrainSet = LearningSet.iloc[;13,:]
TestSet = LearningSet.iloc[13:,:]
```

• The file contains both features & labels, separate them into 4 arrays:

```
Xtrain = TrainSet.drop([`P_NAME', `P_HABITABLE'], axis = 1)
XTest = TestSet.drop([`P_NAME', `P_HABITABLE'], axis = 1)
ytrain = TrainSet.P_HABITABLE
ytest = TestSet.P_HABITABLE
```

 Import the Decision Tree Classifier from sklearn and build the decision tree using the "fit" method.

```
model = DecisionTreeClassifier (random_state = 3)
model.fit(Xtrain, ytrain)
```

#### LET'S SEE WHAT SKLEARN SAYS



#### AND VISUALIZE THE CRITERIA THAT WE FOUND!



Can you figure out the accuracy on the test set?

#### NOTE (AND WE'LL SEE CODE FOR IT): IF YOU USE THE LAST I3 ROWS FOR TRAINING AND THE FIRST 5 FOR TESTING, YOU GET THIS TREE:



and 100% accuracy on test set

Morale: Different train/test split might give significantly different performances, especially when data sets are small.

### KNN ALGORITHM (K NEAREST NEIGHBORS)

Simple, yet powerful!

Only one parameter: k (a small integer)

To make a prediction for a new object, find k closest examples in training set

For classification problems, output the majority class

For regression problems, output the mean of the target property

#### LET'S USE OUR OLD EXAMPLE. WHAT VALUE SHOULD WE USE FOR K?

odd is better!



#### TWEAKABLES!

I) Choose neighborhood radius instead of k

2) Weigh different objects according to distance (inversedistance weighing)

•Any insights on the effects of 1) and 2)?



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I) Choose neighborhood radius instead of k

2) Weigh different objects according to distance (inversedistance weighing)

•Any insights on the effects of I) and 2)?

• If the data have non-uniform density, different choices will have different effects! Needs to be chosen via cross-validation



## **Ensemble Methods**

Two heads are better than one, or「三個臭皮匠,勝過一個諸葛亮」

### Why Ensembles?

- **Statistical:** Multiple minima with same performance (training set too small). Choosing average reduces risk of wrong hypothesis choice.
- **Computational:** get stuck in local minima; results (e.g. decision tree structure + classification) vary strongly depending on training set.
- **Representational:** more expressive than single predictor, e.g.,



Aggregating different linear hypotheses



Linear perceptron hypothesis

### **Ensemble Methods**

 Consider binary classification. Suppose we have N classifiers, each with accuracy p. If the models are independent of each other, the probability that k classifiers are correct:

$$P(k, N, p) = \frac{N!}{k!(N-k)!} p^k (1-p)^{N-k}$$

• Say the ensemble model classifies by a majority vote, i.e.,  $k \ge N/2$ 

$$p(\text{combined model}) = \sum_{i=N/2}^{N} \frac{N!}{i!(N-i)!} p^{i}(1-p)^{N-i}$$

- p(combined model) > p if p > 0.5. For example, if N = 11, p = 0.6, you are encouraged to check that p(combined model) = 0.75.
- Ensemble methods include **Bagging**, **Boosting**, and **Random Forest**.